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Search for chargino and neutralino production using the OPAL detector at $\sqrt{s} = 130-136$ GeV at LEP

OPAL Collaboration

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Abstract

A search for charginos and neutralinos, predicted by supersymmetric theories, has been performed using a data sample of 2.6 pb⁻¹ at a centre-of-mass energy of $\sqrt{s} = 130$ GeV and 2.6 pb⁻¹ at 136 GeV collected with the OPAL detector at LEP during November 1995. No candidate events were observed. The 95% C.L. lower limit on the lightest chargino mass in the Minimal Supersymmetric Standard Model is 65.4 GeV if the universal scalar mass m_0 is greater than 1 TeV, and 58.7 GeV for the smallest m_0 compatible with slepton and sneutrino mass limits obtained at centre-of-mass energies near the Z peak. These limits were obtained under the conditions that the lightest chargino is heavier than the lightest neutralino by more than 10 GeV and tan β is larger than 1.5. The results of a model independent search for charginos and neutralinos are also given.

1. Introduction

In November 1995 the LEP e^+e^- collider at CERN was run for the first time well above the Z peak at centre-of-mass energies (\sqrt{s}) of 130–140 GeV (LEP1.5). This provided an opportunity to search for new particles at these higher energies, but below the $e^+e^- \rightarrow W^+W^-$ threshold.

It is particularly interesting to search for charginos and neutralinos predicted by supersymmetric (SUSY) theories [1]. Charginos $\tilde{\chi}_j^{\pm}$ are the two mass eigenstates formed by the mixing of the fermionic partners of the charged gauge bosons (winos) and those of the charged Higgs bosons. Fermionic partners of the γ (photinos), the Z boson (zinos), and the neutral Higgs bosons mix into the four mass eigenstates called neutralinos $\tilde{\chi}_i^{0.6}$ The lightest neutralino $\tilde{\chi}_1^{0}$ is assumed to be the lightest supersymmetric particle and is therefore stable and invisible if R-parity [2] is conserved. We used the Minimal Supersymmetric Standard Model (MSSM) [3] to guide the analysis but more general cases were also studied.

If charginos are light enough, they can be pairproduced in e^+e^- collisions through γ or Z exchange in the *s*-channel and through sneutrino ($\tilde{\nu}$) exchange in the *t*-channel. The production cross section is fairly large unless the sneutrino is light in which case destructive interference may occur between the *s*-channel and *t*-channel diagrams [4.5].

The details of chargino decay depend on the parameters of the mixing and the masses of the scalar partners of the ordinary fermions [4,5]. The lightest chargino $\tilde{\chi}_1^+$ can decay into a neutralino $\tilde{\chi}_1^0$ and a lepton pair: $\tilde{\chi}_1^+ \rightarrow \tilde{\chi}_1^0 \ell^+ \nu$, or a neutralino and a quark pair: $\tilde{\chi}_1^+ \rightarrow \tilde{\chi}_1^0 q \bar{q}'$ through virtual W, slepton $(\tilde{\ell})$, or scalar quark (\tilde{q}) emission. The effects of the latter

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⁶ The indices i and j are ordered by increasing mass.

are ignored with the assumption that scalar quarks are very heavy. $\tilde{\chi}_1^+$ decay via virtual W emission is dominant in most of the MSSM parameter space; however, if the $\tilde{\chi}_1^+$ is almost a pure wino and the $\tilde{\chi}_1^0$ is almost a pure SUSY partner of the U(1) gauge boson, the $\tilde{\chi}_1^+$ decay via the W boson is suppressed and the dominant decay mode may be $\tilde{\chi}_1^0 \ell^+ \nu$ via a virtual $\tilde{\ell}$ or $\tilde{\nu}$. Due to the energy and momentum carried away by the invisible $\tilde{\chi}_1^0$, chargino events are characterised by large missing energy and transverse momentum imbalance. At $\sqrt{s} = 130-136$ GeV, which is far below the threshold for pair production of real W's or Z's, the backgrounds to chargino pair production are expected to be relatively small.

Previous searches at LEP 1, running at centre-ofmass energies near the Z peak, set lower limits on the mass of the charginos at around $M_Z/2$ using a combination of direct searches and Z width measurements [6]. Similar but more model-dependent limits were obtained by the CDF and D0 Collaborations at the Tevatron [7].

Neutralino pairs $(\tilde{\chi}_i^0 \tilde{\chi}_j^0)$ can be produced through an *s*-channel virtual Z boson or by *t*-channel selectron (\tilde{e}) exchange. At $\sqrt{s} = 130-136$ GeV, the second lightest neutralino $\tilde{\chi}_2^0$ produced in conjunction with a $\tilde{\chi}_1^0$ could give the first direct signal for neutralinos, since single photon events from $e^+e^- \rightarrow \tilde{\chi}_1^0 \tilde{\chi}_1^0 \gamma$ suffer from background from $e^+e^- \rightarrow \nu \bar{\nu} \gamma$. At LEP 1 limits have been obtained from measurements of the width of the Z boson and direct searches [8].

If the $\tilde{\chi}_2^0$ is the lightest visible SUSY particle, it would decay into $\tilde{\chi}_1^0 \ell^+ \ell^-$ or $\tilde{\chi}_1^0 q \bar{q}$ via a Z^{*}, a Higgs boson, a scalar lepton, or a scalar quark. This leads to a similar experimental topology to that for chargino events.

In this paper we report on a direct search for charginos and neutralinos using the data collected with the OPAL detector at $\sqrt{s} = 130-136$ GeV. The data sample collected at 140 GeV was very small and therefore was not used.

2. The OPAL detector and event simulation

2.1. The OPAL detector

The OPAL detector, which is described in detail in [9], is a multipurpose apparatus having nearly

complete solid angle coverage. The central detector consists of a system of tracking chambers providing charged particle tracking over 96% of the full solid angle⁷ inside a 0.435 T solenoidal magnetic field. A lead-glass electromagnetic (EM) calorimeter located outside the magnet coil covers the full azimuthal range with excellent hermeticity in the polar angle range $|\cos \theta| < 0.82$ for the barrel region and $0.81 < |\cos \theta| < 0.984$ for the endcap region. The magnet return yoke is instrumented for hadron calorimetry (HCAL) consisting of barrel and endcap sections along with pole tips that together cover the region $|\cos \theta| < 0.99$. Calorimeters close to the beam axis measure the luminosity using small angle Bhabha scattering events and complete the geometrical acceptance down to 26 mrad. These include the forward detectors (FD) which are lead-scintillator sandwich calorimeters and at smaller angles, silicon tungsten calorimeters (SW) [10] located on both sides of the interaction point. The gap between the endcap EM calorimeter and FD is filled by an additional electromagnetic calorimeter, called the gamma-catcher.

2.2. Monte Carlo event simulation

The Monte Carlo generator SUSYGEN [11] was used to produce chargino and neutralino pair events. We also used a Monte Carlo generator similarly based on the calculation of the differential cross sections by Bartl et al. [4] for chargino pair production. The cross sections calculated by both generators agree very well. In both generators initial state radiation was included, and the JETSET7.4 package [12] was used for the hadronization of the quark-antiquark system in the chargino or neutralino hadronic decay with parameters specified in Ref. [13].

The most important parameters governing the chargino detection efficiency are the mass of the chargino $m_{\tilde{\chi}_1^+}$ and the mass difference between the chargino and the lightest neutralino: $\Delta M_+ \equiv (m_{\tilde{\chi}_1^+} - m_{\tilde{\chi}_1^0})$. For the case of the neutralino, $m_{\tilde{\chi}_2^0}$ and $\Delta M_0 \equiv (m_{\tilde{\chi}_2^0} - m_{\tilde{\chi}_1^0})$ mainly determine the efficiency. The other important parameters are the decay branch-

⁷ A right-handed coordinate system is adopted, where the x-axis points to the centre of the LEP ring, and positive z is along the electron beam direction. The angles θ and ϕ are the polar and azimuthal angles, respectively.

ing fractions. We generated $\tilde{\chi}_1^+ \tilde{\chi}_1^-$ events at 30 points in the $m_{\tilde{\chi}_1^+} - \Delta M_+$ plane. For the $\tilde{\chi}_1^0 \tilde{\chi}_2^0$ process, events were generated at 42 points in the $m_{\tilde{\chi}_2^0} - \Delta M_0$ plane.

There are several sources of background to the chargino and neutralino signals:

- Two photon processes are the most important background for the case of small ΔM_+ (or ΔM_0), since signal events have small visible energy and small transverse momentum relative to the beam direction. We used the Monte Carlo generator PYTHIA [12] for generating events from two photon processes where the virtual mass Q^2 of both photons is smaller than 1.3 GeV^2 and the invariant mass of the photonphoton system $(M_{\gamma\gamma})$ is greater than 3 GeV. For higher Q^2 events, the generator TWOGEN [14] was used. Event samples for all the possible processes (final state hadrons from point-like $\gamma \gamma \rightarrow q\bar{q}$ processes and from vector meson dominance, and all $e^+e^-\ell^+\ell^$ final states) were generated. Two photon events were not generated in the region $Q^2 < 1.3 \text{ GeV}^2$ and $M_{\gamma\gamma} < 3$ GeV. This region did not represent a serious background to the search presented here.

- τ pairs are the dominant background for the monojet signal when one of the τ 's decays into a very low momentum electron and two energetic neutrinos. $\tau^+\tau^-\gamma$ events are also a potential source of background for the topology of two acoplanar jets. The event generator KORALZ [15] was used for the generation of $\tau^-\tau^-(\gamma)$ events.

- Ordinary multijet hadronic events in which one or more jet momenta are mismeasured are the dominant background for the large mass difference case. The PYTHIA generator [12] was used to simulate hadronic events.

- Finally, four fermion processes in which at least one of the fermions is a neutrino constitute a serious background. Although the cross sections for these processes are small at \sqrt{s} below the W⁺W⁻ or ZZ threshold, their event topology is similar to that of the signal. Since the interference effects of many diagrams are important below the W⁺W⁻ threshold, we used an event generator based on helicity amplitude calculations, which take into account all the relevant diagrams and interference effects [16]. The package also includes initial state photon radiation.

Generated signal and background events were processed through the full simulation of the OPAL detector [17], and the same event analysis chain was applied to simulated events as to the data.

3. Analysis

The present analysis is based on the data collected during the 1995 November run of LEP. Data used in this analysis corresponded to an integrated luminosity of 2.6 pb⁻¹ at a centre-of-mass energy of \sqrt{s} = 130 GeV and 2.6 pb⁻¹ at 136 GeV⁸.

The experimental signatures for $\tilde{\chi}_1^+ \tilde{\chi}_1^-$ or $\tilde{\chi}_1^0 \tilde{\chi}_2^0$ events are an acoplanar pair of particles or jets, or a monojet topology with large transverse momentum with respect to the beam axis. If the mass difference between $\tilde{\chi}_1^+$ and $\tilde{\chi}_1^0$ is as small as 5 GeV, the visible energy as well as the transverse momentum becomes small. The selection was designed to maintain reasonable efficiency for these cases.

Good charged tracks were selected using the same track quality requirements as in Ref. [18]. Tracks were required to have at least 20 measured space points, more than 50% of the hits geometrically expected, and transverse momentum exceeding 100 MeV. Electromagnetic clusters in the barrel region were required to have an energy of at least 170 MeV, and the clusters in the endcaps to have an energy of at least 250 MeV and contain at least two adjacent lead glass blocks. Clusters in the hadron calorimeters were required to have an energy of at least 0.6 GeV in the barrel and endcaps, and at least 2 GeV in the pole tips.

The measurements of visible energy, mass and total transverse momentum of the events were performed by the method used in Ref. [18]. Four momentum vectors were formed for each track and calorimeter cluster, and then summed. The calorimeter clusters were treated as massless particles. In order to reduce the effects of double counting, four vectors based on the average expected energy deposition in the calorimeters for each charged track were then subtracted. This average energy deposition was parametrized as a function of track momentum and polar angle. The procedure used avoids the need for detailed matching between tracks and calorimeter clusters.

 $^{^{8}}$ The actual centre-of-mass energies were estimated to be 130.26 and 136.23 GeV, but for the purpose of this analysis the nominal values of 130 and 136 GeV were assumed.

When calculating angles, such as the polar angle of the thrust axis, θ_{thrust} , or the acoplanarity angle (defined below), the charged track momenta are more important than the calorimeter information due to their better angular resolution. The track momenta and momentum vectors of the EM or hadron calorimeter clusters not associated with charged tracks were first summed. In the case where a calorimeter cluster had associated good charged tracks, the scalar sum of the associated charged track momenta was subtracted from the cluster energy to reduce double counting. If the energy of a cluster was smaller than the scalar sum of the momenta of the associated tracks, the cluster energy was not used. If the number of charged particles in the event was equal to two, this method was also used to calculate all the variables including visible energy and invariant masses. These events are mainly lepton pairs while the algorithm of Ref. [18] was optimised for hadronic jets.

To select chargino and neutralino candidates, the number of good charged tracks was required to be at least two, and the ratio of the number of good tracks to the total number of tracks to be greater than 0.2.

To reduce background from two photon processes and from multihadronic events where a jet axis is close to the beam direction, the total energy deposited in each silicon tungsten calorimeter had to be less than 5 GeV, less than 2 GeV in each forward calorimeter, and less than 5 GeV in each side of the gammacatcher. In addition, the visible energy in the region of $|\cos \theta| > 0.8$ should be less than 30% of the total visible energy (cut 1). After this cut the distribution of total visible energy normalised by \sqrt{s} is shown in Fig. 1a and the distribution of event transverse momentum (P_t , magnitude of the vector sum of transverse momentum components) measured without including the hadron calorimeter is shown in Fig. 1b. Since two photon Monte Carlo events were generated with $M_{\gamma\gamma} > 3$ GeV, the visible mass of the events was required to be larger than 3 GeV just for these plots in order to demonstrate that the normalization of the two photon events agrees with the Monte Carlo prediction. The events with small invariant mass were eliminated from the data by demanding that the P_t be greater than 4 GeV and the event transverse momentum measured including the hadron calorimeter (P_t^{HCAL}) should be greater than 5 GeV (cut 2). Although most of the events from two photon processes were rejected by the

 P_t cut, the P_t^{HCAL} cut was applied to reject occasional two photon events with a high transverse momentum neutral hadron.

"Radiative return" events from $e^+e^- \rightarrow Z\gamma$, where the γ escaped down the beam pipe, were rejected by requiring that the polar angle of the missing momentum direction θ_{miss} satisfy $|\cos \theta_{\text{miss}}| < 0.7$ (cut 3). Events with a photon of energy greater than 15 GeV and accompanied by no good tracks (except for track pairs identified as photon conversions) within a cone of half angle 25° from the photon direction were rejected (cut 4).

To remove a large fraction of the multihadron events the visible energy was required to be less than $0.7\sqrt{s}$ (cut 5). Finally, events that were kinematically consistent with $\tau^+\tau^-(\gamma)$ were rejected (cut 6). This $\tau^+\tau^-(\gamma)$ rejection was accomplished by requirements on multiplicity, momentum sum and invariant mass of charged particles in each hemisphere of the event, and a requirement on the acoplanarity angle which depended on the absolute value and the polar angle of the charged particle momentum sum in each hemisphere.

The tracks and the clusters in an event were then divided into two hemispheres defined by the plane perpendicular to the thrust axis. If one of the hemispheres had an energy of less than 1 GeV and also contained no good tracks, the event was categorised as a monojet event. Otherwise the event was classified as a dijet event. According to these criteria, ten events in the data were classified as dijet events, and four as monojet events.

For dijet events, the additional selection criteria below were applied. The polar angle of the event thrust axis was required to satisfy $|\cos\theta_{\text{thrust}}| < 0.9$ (cut A1). No events were removed by this requirement, but it ensures good resolution of the acoplanarity angle. The acoplanarity angle ϕ_{acop} was defined as $\pi - \phi_{\text{open}}$, where ϕ_{open} is the azimuthal opening angle between the directions of the two momentum sums of the particles in the thrust hemispheres. The acoplanarity angle (ϕ_{acop}) between the momentum sums in the two hemispheres was required to be greater than 20° (cut A2). The acoplanarity angle distribution just before the cut is shown in Fig. 1c. The expected distribution for $\tilde{\chi}_1^+ \tilde{\chi}_1^-$ with $m_{\tilde{\chi}_1^+} = 60$ GeV and $m_{\tilde{\chi}_1^0} = 30$ GeV is also shown.

For events classified as monojets, additional se-



Fig. 1. (a) The distribution of the total visible energy normalised to \sqrt{s} and (b) the distribution of P_t after cut (1). (c) The acoplanarity angle distribution after cut (A1). The prediction for $\tilde{\chi}_1^+ \tilde{\chi}_1^-$ events is also shown for $m_{\tilde{\chi}_1^+} \approx 60$ GeV and $m_{\tilde{\chi}_1^0} = 30$ GeV with a production cross section of 20 pb as a dashed histogram. (d) The visible mass distribution for monojet events. The prediction for $\tilde{\chi}_1^0 \tilde{\chi}_2^0$ events is also shown for $m_{\tilde{\chi}_2^0} = 80$ GeV and $m_{\tilde{\chi}_1^0} = 30$ GeV with a production cross section of 5 pb as a dashed histogram. In all the plots, the hatched area indicates the prediction for $q\bar{q}(\gamma)$ events, the grey area for $\ell^+\ell^-(\gamma)$ events, the open area for two photon processes, and the double hatched area for four fermion processes ($\nu \bar{\nu} \ell^+ \ell^-$, $\nu \bar{\nu} q\bar{q}$, $\nu^\ell q\bar{q}'$, $\tau^+ \tau^- q^+, \tau^- \tau^+ \tau^-$ and $\tau^+ \tau^- \mu^+ \mu^-$).

lection criteria were required. To reject $e^+e^- \rightarrow e^+e^-\tau^+\tau^-$ events where one of the τ 's decays into a soft electron and to reject four fermion processes $e^+e^- \rightarrow Z^*\gamma^* \rightarrow \nu\bar{\nu}f\bar{f}$, where f is a charged lepton or a quark, the visible mass of the event was required to be greater than 2 GeV (cut B1). The visible mass distribution just before the cut is shown in Fig. 1d. The expected distribution for $\tilde{\chi}_1^0 \tilde{\chi}_2^0$ with $m_{\tilde{\chi}_2^0} = 80$ GeV and $m_{\tilde{\chi}_1^0} = 30$ GeV is also shown. Furthermore, the ratio of the visible mass to the visible energy ($M_{\rm vis}/E_{\rm vis}$) was required to be larger than 0.4 (cut B2).

No events were observed in the data after the above cuts.

The remaining numbers of events after each cut are listed in Table 1. For comparison the table also shows the corresponding numbers of simulated events for background processes and for three samples of simulated $\tilde{\chi}_1^+ \tilde{\chi}_1^-$ and $\tilde{\chi}_1^0 \tilde{\chi}_2^0$ events.

The efficiency for $\tilde{\chi}_1^+ \tilde{\chi}_1^-$ events is about 42% for $m_{\tilde{\chi}_1^+} = 60$ GeV and $m_{\tilde{\chi}_1^0} = 30$ GeV at $\sqrt{s} = 130$ GeV for a $\tilde{\chi}^+$ decay via a virtual W. The efficiency de-

Table 1

The remaining numbers of events normalised to the integrated luminosity of the data for various background processes are compared with data after each cut. Numbers for three simulated event samples of $\tilde{\chi}_1^+ \tilde{\chi}_1^-$ with $\tilde{\chi}_1^+ \to \tilde{\chi}_1^0 W^*$, and $\tilde{\chi}_2^0 \tilde{\chi}_2^0$ with $\tilde{\chi}_2^0 \to \tilde{\chi}_1^0 Z^*$ are also given (starting from 1000 events for each). The numbers given after cuts (A1) or (A2) are for dijet events only, and the numbers after cuts (B1) or (B2) are for monojet events only. The numbers of events expected from two photon processes do not include the region $M_{\gamma\gamma} < 3$ GeV with $Q^2 < 1.3$ GeV².

	Data	Total bkg.	$q \bar{q}(\gamma)$	$ au au(\gamma)$	'γγ'	4-f	$\tilde{\chi}_1^+ \tilde{\chi}_1^-$	$\tilde{\chi}_1^+ \tilde{\chi}_1^-$	$ ilde{\chi}^0_1 ilde{\chi}^0_2$	
$m_{\tilde{\chi}^{\pm}}$ (GeV)							60	60	-	-
$m_{\tilde{\chi}^0}$ (GeV)							55	30	30	
$m_{\tilde{\chi}_2^0}^{\Lambda_1}$ (GeV)							-	-	80	
no cuts	_	-	1650	115.3	105k	11.1	1000	1000	1000	
cut (1)	37910	8896	797.9	58.6	7104	2.8	680	733	758	
cut (2)	511	546.2	352.3	44.8	8.0	2.5	180	677	746	
cut (3)	207	245.8	162.5	30.5	3.1	2.0	176	556	632	
cut (4)	169	206.3	132.6	26.7	3.1	2.0	176	556	632	
cut (5)	35	34.5	6.8	22.3	2.9	1.9	176	556	631	
cut (6)	14	16.0	6.8	4.2	2.9	1.9	176	554	631	
dijet	10	13.7	6.8	4.2	2.1	0.42	104	483	233	
cut (A1)	10	13.7	6.8	4.2	2.1	0.40	103	474	233	
cut (A2)	0	0.75	0.0	0.17	0.26	0.33	91	366	191	
monojet	4	2.31	0.0	0.03	0.78	1.50	72	71	398	
cut (B1)	0	0.88	0.0	0.02	0.00	0.85	70	71	382	
cut (B2)	0	0.18	0.0	0.01	0.00	0.17	65	52	90	
(A2+B2)	0	0.93	0.0	0.18	0.26	0,49	156	418	281	

creases to 15% for $m_{\tilde{\chi}_1^0} = 55$ GeV. For the case of purely leptonic decays, the efficiencies are about 5% lower than for the decay via W^{*} for the same $\tilde{\chi}_1^+$ and $\tilde{\chi}_1^0$ mass combination. The efficiency for $\tilde{\chi}_1^0 \tilde{\chi}_2^0$ events is typically 30-40% over a wide range of $\tilde{\chi}_2^0$ and $\tilde{\chi}_1^0$ masses. The efficiency drops to 10% for the small mass difference of $\Delta M_0 = 5$ GeV. For the extraction of limits described in the following section the efficiency at an arbitrary point of $m_{\tilde{\chi}_1^+} (m_{\tilde{\chi}_2^0})$ and $\Delta M_+ (\Delta M_0)$ was interpolated using a polynomial fit to the efficiencies determined using Monte Carlo simulations.

4. Results

No evidence for chargino or neutralino production is observed in the data; therefore, limits are calculated. In the first analysis, upper limits on the cross sections as functions of chargino and neutralino masses were calculated for 100% branching fractions into specific decay modes (model independent analysis). In the second analysis, limits on chargino and neutralino masses were calculated within the framework of the MSSM, and limits in the MSSM parameter space were obtained.

4.1. Systematic errors

The common systematic errors for the model independent analysis and the analysis within the framework of the MSSM were first considered. The systematic error on the integrated luminosity was 1.5%, which was evaluated from half the difference between the luminosities measured by the FD detector and the SW detector. The relative errors on the signal efficiencies due to Monte Carlo statistics and the interpolation errors at an arbitrary point of $m_{\tilde{\chi}_1^+}$ ($m_{\tilde{\chi}_2^0}$) and $m_{\tilde{\chi}_1^0}$ were evaluated to be 2–10% depending on the combination of the masses and the branching fractions. The systematic error due to the trigger efficiency was negligible for the selected signal events. The various systematic errors were summed quadratically and subtracted from the signal efficiency.

4.2. Limits on the production cross sections

We obtain upper limits on the production cross sections at the 95% confidence level (C.L.) for $\tilde{\chi}_1^+ \tilde{\chi}_1^$ and $\tilde{\chi}_1^0 \tilde{\chi}_2^0$ assuming specific decay modes. These limits do not depend on the details of the SUSY models. The contours of the upper limits for the $\tilde{\chi}_1^+ \tilde{\chi}_1^$ cross sections at $\sqrt{s} = 136$ GeV are shown in Fig. 2, assuming $\tilde{\chi}_1^+ \rightarrow \tilde{\chi}_1^0 W^*$ with 100% branching fraction (Fig. 2a) or $\tilde{\chi}_1^+ \to \tilde{\chi}_1^0 \ell^+ \nu$ via $\tilde{\ell}$ or $\tilde{\nu}$ with 100% branching fraction (Fig. 2b). The cross sections at \sqrt{s} = 130 GeV were assumed to be the same as at 136 GeV except if $2m_{\tilde{\chi}_1^+} > 130$ GeV. Similarly, the contours of the upper limit for the $\tilde{\chi}_1^0 \tilde{\chi}_2^0$ cross sections are shown for $\tilde{\chi}_2^0 \rightarrow \tilde{\chi}_1^0 Z^*$ with 100% branching fraction (Fig. 2c), or $\tilde{\chi}_2^0 \to \tilde{\chi}_1^0 \ell^+ \ell^-$ via $\tilde{\ell}$ with 100% decay branching fraction (Fig. 2d). The Standard Model branching fractions were used for the W* and Z* decays including the invisible decay mode $Z^* \rightarrow \nu \tilde{\nu}$.

If the cross section for $\tilde{\chi}_1^+ \tilde{\chi}_1^-$ is larger than 3.5 pb, we can exclude the $\tilde{\chi}_1^+$ for masses up to 65.0 GeV at the 95% C.L. for $\Delta M_+ \geq 10$ GeV independent of the decay mode, and almost up to the kinematical boundary for $\tilde{\chi}_1^+$ decay via W*. From Fig. 2c, we can also exclude the $\tilde{\chi}_2^0$ for masses almost up to the kinematical boundary of $(m_{\tilde{\chi}_1^0} + m_{\tilde{\chi}_2^0}) < \sqrt{s}$ at the 95% C.L. for $\Delta M_0 \geq 13$ GeV, if the cross section for $\tilde{\chi}_1^0 \tilde{\chi}_2^0$ with $\tilde{\chi}_2^0 \to \tilde{\chi}_1^0 Z^*$ is larger than 6 pb. The region $(m_{\tilde{\chi}_1^0} + m_{\tilde{\chi}_2^0}) < M_Z$, which is accessible at LEP 1, is not considered in this analysis. For $\tilde{\chi}_1^0 \tilde{\chi}_2^0$ produced through Z boson decay better limits were obtained at LEP 1 [8].

4.3. Limits on the minimal supersymmetric standard model

The negative results of the above searches can be interpreted in the framework of the MSSM where the gaugino-higgsino sector of the theory is completely determined by three parameters: M_2 , the SU(2) gaugino mass at the weak scale; μ , the mass coupling strength between the two Higgs superfields; and tan β , the ratio $\langle v_2 \rangle / \langle v_1 \rangle$ of the vacuum expectation values of the two Higgs doublets. For a given set of these parameters there are, at tree level, unique relations [11] that determine masses and coupling constants of all gauginos. The scanned regions of the parameters were $0 \le M_2 \le 1500$ GeV, $-400 \le \mu \le 400$ GeV, for two values of tan β : tan $\beta = 1.5$, a small value interesting for infrared fixed-point predictions for the top quark [19], and tan $\beta = 35$, approximately equal to the ratio of the top quark mass to the b quark mass favoured by Yukawa coupling unification at a large mass scale [20]. A more general case of all values of tan $\beta > 1$ was also considered. The scanned ranges of M_2 and μ were checked to be large enough so that the exclusion regions presented change negligibly for larger ranges.

The f spectrum is further constrained by assuming soft symmetry breaking at the Grand Unification scale (GUT) with a common mass scale m_0 (defined at the GUT scale) [21]. A light m_0 sets the masses of the $\tilde{\nu}$ and $\tilde{\ell}$ to low values, thereby enhancing *t*-channel exchange diagrams that may have destructive interference with s-channel diagrams reducing the cross section for chargino production. Small values of m_0 also tend to enhance the leptonic branching ratio of charginos. On the other hand, fixing m_0 to high values decouples the $\tilde{\nu}$ and \tilde{e} from the theory, thereby enhancing the chargino production rate. We therefore present results here in the two scenarios: $m_0 = 1$ TeV, and the other extreme, the smallest m_0 consistent with light $\tilde{\ell}$ and $\tilde{\nu}$ not yet excluded by LEP 1, namely $m_{\tilde{\ell}} > 1$ 45.0 GeV and $m_{\tilde{\nu}} > 41.8$ GeV [22].

In the MSSM, production cross sections and decay branching fractions of charginos and neutralinos are determined by the four SUSY parameters; M_2 , μ , tan β and m_0 . In regions where chargino production is suppressed, SUSY signatures can be explored via neutralino production. The analysis presented here is therefore based on combining both the neutralino and chargino searches into one analysis. At each point in the MSSM parameter space the number of events expected for $\tilde{\chi}_1^+ \tilde{\chi}_1^-$, $\tilde{\chi}_1^0 \tilde{\chi}_2^0$, as well as $\tilde{\chi}_2^0 \tilde{\chi}_2^0$ were added. The decay branching fractions were calculated and the corresponding detection efficiencies were used in order to set exclusion limits at the 95% C.L. in three different MSSM parameter planes. Since $\tilde{\chi}_2^0$ can decay into $\tilde{\chi}_{1}^{0}$ and the lightest Higgs boson, the efficiency for this case was checked and found to be similar to that for the $\tilde{\chi}_2^0 \rightarrow \tilde{\chi}_1^0 q \tilde{q}$ decay via a Z^{*}.

The $\tilde{\chi}_2^0$ can decay into $\nu \bar{\nu} \tilde{\chi}_1^0$ via a Z^{*} or $\tilde{\chi}_1^0 \tilde{\chi}_1^0 \tilde{\chi}_1^0$



Fig. 2. The 95% C.L. upper limits for the production cross sections at $\sqrt{s} = 136 \text{ GeV}$ for $\tilde{\chi}_1^+ \tilde{\chi}_1^-$ assuming (a) $\tilde{\chi}_1^+ \rightarrow \tilde{\chi}_1^0 W^{*+}$ with 100% branching fraction or (b) $\chi_1 \rightarrow \tilde{\chi}_1^0 \ell^+ \nu$ via $\tilde{\ell}$ or $\tilde{\nu}$ with 100% branching fraction. The contours of the upper limits for the $\tilde{\chi}_1^0 \tilde{\chi}_2^0$ cross sections at $\sqrt{s} = 136 \text{ GeV}$ are shown for (c) $\tilde{\chi}_2^0 \rightarrow \tilde{\chi}_1^0 Z^*$ with 100% branching fraction, or (d) $\tilde{\chi}_2^0 \rightarrow \tilde{\chi}_1^0 \ell^+ \ell^-$ via $\tilde{\ell}$ with 100% decay branching fraction. The regions $(m_{\tilde{\chi}_1^0} + m_{\tilde{\chi}_2^0}) < M_Z$, which are accessible at LEP 1, are not considered in this analysis and are indicated in the plots. The kinematical boundaries for production and decay at $\sqrt{s} = 136 \text{ GeV}$ are shown by dashed lines.

via the lightest Higgs, leading to invisible events. In most of the MSSM parameter space the $\tilde{\chi}_1^+$ is lighter than the $\tilde{\chi}_2^0$. Therefore the cascade decay of $\tilde{\chi}_2^0 \rightarrow \tilde{\chi}_1^{\pm} q \bar{q}' \text{ or } \rightarrow \tilde{\chi}_1^{\pm} \ell^{\mp} \nu$ with subsequent decay of the $\tilde{\chi}_1^{\pm}$ is possible. These effects were taken into account in calculating the limits.

The results are shown in Figs. 3 and 4. Fig. 3 shows the exclusion contours in the M_2 - μ plane for the case $m_0 = 1$ TeV (Fig. 3a) and a light m_0 (Fig. 3b) for tan $\beta = 1.5$. The same plots are shown for tan $\beta = 35$ (Fig. 3c and 3d). Most of the excluded regions are due to the limits from the $\tilde{\chi}_1^+ \tilde{\chi}_1^-$ search. The $\tilde{\chi}_1^0 \tilde{\chi}_2^0$ signal contributes only in the limited area near the kinematical boundary of the large M_2 region. Fig. 4 shows the region excluded at 95% C.L. in the $m_{\tilde{\chi}_1^0} - m_{\tilde{\chi}_1^1}$ plane. The limits are shown for the combinations of two m_0 values and two tan β values. Similar exclusion regions in the $m_{\tilde{\chi}_1^0} - m_{\tilde{\chi}_2^0}$ plane are also shown in Fig. 4. Again, most of the excluded regions are due to the limits from the $\tilde{\chi}_1^+ \tilde{\chi}_1^-$ search.

Fig. 4e and 4f show the regions excluded at 95% C.L. in the $m_{\tilde{\chi}_1^0} - m_{\tilde{\chi}_1^+}$ plane assuming $m_0 = 1$ TeV or the minimum m_0 . The exclusion regions were found by scanning through $\tan \beta$ ($\tan \beta > 1$) to find the smallest number of events expected (sum of $\tilde{\chi}_1^+ \tilde{\chi}_1^-$, $\tilde{\chi}_1^0 \tilde{\chi}_2^0$ and $\tilde{\chi}_2^0 \tilde{\chi}_2^0$ events after all the cuts) at any point of the M_2 - μ plane and for any value of tan β considered.

From the above plots one can set lower limits at 95%



Fig. 3. The exclusion contours in the M_2 - μ plane for the case (a) $m_0 = 1$ TeV and (b) minimum m_0 with tan $\beta = 1.5$. The same plots are shown for tan $\beta = 35$ (c and d). The minimum m_0 is defined to be the smallest m_0 possible to comply with the limits on $\tilde{\ell}$ and $\tilde{\nu}$ at LEP 1. The region excluded by the analysis of LEP 1 data is also shown. The kinematical boundary at $\sqrt{s} = 136$ GeV for $\tilde{\chi}_1^+ \tilde{\chi}_1^-$ production is shown by dashed lines and for $\tilde{\chi}_1^0 \tilde{\chi}_2^0$ production as dotted lines.

C.L. on the chargino and neutralino masses as given in Table 2. Definitive lower limits for neutralino masses are not possible when considering the entire region $\tan \beta > 1$ since for values of $\tan \beta$ close to 1, typical considered mass values of $\tilde{\chi}_1^0$ and $\tilde{\chi}_2^0$ can result from MSSM parameters that give chargino masses that are not kinematically accessible for $\sqrt{s} = 130-136$ GeV.

5. Summary and conclusion

We have analyzed a data sample corresponding to an integrated luminosity of 2.6 pb⁻¹ at $\sqrt{s} = 130$ GeV and 2.6 pb⁻¹ at 136 GeV collected with the OPAL detector to search for pair production of charginos and neutralinos predicted by supersymmetric theories. No events remained after the selection cuts. This is consistent with the expected background of 0.9 events. Within the framework of MSSM and for the conditions $\Delta M_+ \geq 10$ GeV and $\Delta M_0 \geq 10$ GeV one can set lower limits at 95% C.L. on the chargino and neutralino masses: $m_{\tilde{\chi}_1^0} > 21.4$ GeV, $m_{\tilde{\chi}_2^0} > 47.5$ GeV and $m_{\tilde{\chi}_1^+} > 65.4$ GeV for tan $\beta = 1.5$. $m_{\tilde{\chi}_1^0} > 35.2$ GeV, $m_{\tilde{\chi}_2^0} > 67.5$ GeV and $m_{\tilde{\chi}_1^+} > 65.6$ GeV are the limits obtained for tan $\beta = 35$. These limits were obtained for $m_0 = 1$ TeV. For the smallest m_0 possible to comply with the LEP 1 $\tilde{\ell}$ and $\tilde{\nu}$ limits, the mass limits are reduced by at most 7 GeV for the chargino case.



Fig. 4. The 95% C.L. excluded region in the $m_{\tilde{\chi}_1^0} - m_{\tilde{\chi}_1^+}$ plane within the framework of the MSSM for the case of minimum m_0 (light shaded region) and $m_0 = 1$ TeV (extending to dark shaded region) for (a) tan $\beta = 1.5$ and (b) tan $\beta = 35$. The region excluded by the analysis of LEP 1 data is also shown. The thick solid lines bound the region that is physically accessible for the given region of MSSM parameter space. The kinematical boundaries for production and decay at $\sqrt{s} = 136$ GeV are shown by dashed lines. Similar plots for neutralino masses for (c) tan $\beta = 1.5$ and (d) tan $\beta = 35$. The regions excluded at 95% C.L. (e-f) in the $m_{\tilde{\chi}_1^0} - m_{\tilde{\chi}_1^+}$ plane conservatively obtained by scanning through tan β (tan $\beta > 1$) to find the smallest number of events expected (sum of $\tilde{\chi}_1^+ \tilde{\chi}_1^-, \tilde{\chi}_1^0 \tilde{\chi}_2^0$ and $\tilde{\chi}_2^0 \tilde{\chi}_2^0$ events after all the cuts) at any point of the plane and value of tan β for (e) $m_0 = 1$ TeV and (f) minimum m_0 .

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Table 2

Lower limits at 95% C.L. on the chargino mass $m_{\tilde{\chi}_1^+}$, the lightest neutralino mass $m_{\tilde{\chi}_1^0}$, and the second lightest neutralino mass $m_{\tilde{\chi}_2^0}$. These limits are given for the smallest m_0 possible to comply with the LEP 1 $\tilde{\ell}$ and $\tilde{\nu}$ limits and also for $m_0 = 1$ TeV with the mass-difference conditions $\Delta M_+ \geq 10$ GeV and $\Delta M_0 \geq 10$ GeV.

Mass	$\tan \beta = 1.5$		$\tan \beta = 35$	$\tan \beta > 1$		
	Min. m_0	$m_0 = 1$ TeV	Min. m_0	$m_0 = 1 \text{ TeV}$	$m_0 = 1 \text{ TeV}$	
$m_{\tilde{v}_{\perp}^{\perp}}$ (GeV)	> 60.7	> 65.4	> 58.7	> 65.6	> 63.9	
$m_{\tilde{\chi}_{1}^{0}}^{0}$ (GeV)	> 12.0	> 21.4	> 32.0	> 35.2	-	
$m_{\tilde{\chi}_2^0}$ (GeV)	> 45.3	> 47.5	> 61.0	> 67.5	-	

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References

- Y. Gol'fand and E. Likhtam, JETP Lett. 13 (1971) 323;
 D. Volkov and V. Akulov, Phys. Lett. B46 (1973) 109;
 J. Wess and B. Zumino, Nucl. Phys. B70 (1974) 39.
- [2] P. Fayet, Unification of the Fundamental Particle Interactions, Plenum Press (1980), p. 587.
- [3] H. P. Nilles, Phys. Rep. 110 (1984) 1;
- H. E. Haber and G. L. Kane, Phys. Rep. 117 (1985) 75.
- [4] A. Bartl, H. Fraas and W. Majerotto, Z. Phys. C30 (1986) 441;

A. Bartl, H. Fraas and W. Majerotto, Z. Phys. C41 (1988) 475;

A. Bartl, H. Fraas, W. Majerotto and B. Mösslacher, Z. Phys. C55 (1992) 257.

[5] M. Chen, C. Dionisi, M. Martinez and X. Tata, Phys. Rep. 159 (1988) 201;

J. L. Feng and M. J. Strassler, Phys. Rev. D51 (1995) 4661.

[6] L3 Collab., B. Adeva et al., Phys. Lett. B233 (1989) 530;
ALEPH Collab., D. Decamp et al., Phys. Lett. B236 (1990) 86;
OPAL Collab., M.Z. Akrawy et al., Phys. Lett. B240 (1990)

261;

DELPHI Collab., P. Abreu et al., Phys. Lett. B247 (1990) 157;

ALEPH Collab., D. Decamp et al., Phys. Rep. 216 (1992) 253.

- [7] CDF Collab., J. Hauser, Searches for SUSY by CDF, in the Proc. of the 10th Topical Workshop on Proton-Antiproton Collider Physics, Fermilab, IL, USA, May 1995; D0 Collab., S. Abachi et al., Search for $\tilde{W}_1 \tilde{Z}_2$ Production via Trilepton Final States in pp Collisions at $\sqrt{s} = 1.8$ TeV, FERMILAB Pub-95/385-E, submitted to Phys. Rev. Lett.
- [8] OPAL Collab., G. Alexander et al., Topological Search for the Production of Neutralinos and Scalar Particles, CERN-PPE/96-019, submitted to Phys. Lett.
- [9] OPAL Collab., K. Ahmet et al., Nucl. Instr. Meth. A305 (1991) 275.
- [10] B.E. Anderson et al., IEEE Transactions on Nuclear Science 41 (1994) 845.
- S. Katsanevas and S. Melanchroinos, to be published in the LEP200 workshop report, CERN 1995. The formulae found in the following references are employed: Ref. [4] and A. Bartl, H. Fraas, and W. Majerotto, Nucl. Phys. B278 (1986) 1, and Z. Phys. C34 (1987) 411.
- [12] T. Sjöstrand, Comp. Phys. Comm. 39 (1986) 347;
 T. Sjöstrand, PYTHIA 5.7 and JETSET7.4 Manual, CERN-TH 7112/93.
- [13] OPAL Collab., G. Alexander et al., CERN PPE/95-126, to be published in Z. Phys.
- [14] A. Buijs et al., Comp. Phys. Comm. 79 (1994) 523.
- [15] S. Jadach, B. F. L. Ward, Z. Was, Comp. Phys. Comm. 79 (1994) 503.

- [16] The calculation of the cross section is based on the method described in the preprint by H. Murayama, I. Watanabe and K. Hagiwara, KEK Report 91-11 (1991). The initial state radiation is implemented according to a formula in F. A. Berends, R. Pittau and R. Kleiss, Nucl. Phys. B426 (1994) 344.
- [17] J. Allison et al., Nucl. Instr. Meth. A317 (1992) 47.
- [18] OPAL Collab., M.Z. Akrawy et al., Phys. Lett. B253 (1991)
 511. In the present analysis an improved parametrization of the energy measurement was used.
- [19] M. Carena and C.E.M. Wagner, Nucl. Phys. B452 (1995) 45.
- [20] J. Ellis, S. Kelley, and D.V. Nanopoulos, Phys. Lett. B287 (1992) 95.
- [21] M. Drees and M. M. Nojiri, Nucl. Phys. B369 (1992) 54.
 M. Carena, S. Dimopolous, S. Raby and C.E.M. Wagner, Phys. Rev. D52 (1995) 4133.
- [22] Particle Data Group, Phys. Rev. D50 (1994) 1173.