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Title: JOZSO, a computer code for calculating broad neutron resonances in phenomenological nuclear potentials

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Abstract: A renewed version of the computer code GAMOW (T. Vertse, K. F. Pál, Z. Balogh, Computer Physics Communications **27**, 309 (1982).) is given in which the difficulties in calculating broad neutron resonances are amended. New types of phenomenological neutron potentials with strict finite range are built in. Landscape of the S-matrix can be generated on a given domain of the complex wave number plane and S-matrix poles in the domain are localized. Normalized Gamow wave functions and trajectories of given poles can be calculated optionally.

Detailed Response to Reviewers

Answer to the referee.

Thank very much for the referee her/his comments and reading carefully the manuscript.

We included in the paper most of the corrections proposed by the referee.

To the question

"page 14, equation line 11-12: why is the sum limit equal to 50? Is this just a typo?":

It is not a typo actually. We used the limit 50 in order to have a greater accuracy. For example, if $r=0.5$ then the first term of the sum which is less than the machine accuracy is the 50-th term.

In Subsection 2.4 we changed the first few lines in order to be more clear. In Eq. (31) we included a subscript 0 of the variable k .

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9 JOZSO, a computer code for calculating broad neutron
10 resonances in phenomenological nuclear potentials
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23 **Abstract**

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25 A renewed version of the computer code GAMOW [1] is given in which
26 the difficulties in calculating broad neutron resonances are amended. New
27 types of phenomenological neutron potentials with strict finite range are built
28 in. Landscape of the S-matrix can be generated on a given domain of the
29 complex wave number plane and S-matrix poles in the domain are localized.
30 Normalized Gamow wave functions and trajectories of given poles can be
31 calculated optionally. ¹
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34 *Keywords:* resonance, finite range potential,
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37 **PROGRAM SUMMARY**

38 *Program Title:* JOZSO

39 *Licensing provisions:* GPLv3

40 *Programming language:* Fortran 90

41 *Supplementary material:* A readme file: <https://github.com/czylabsonasa/jozso>

42 *Nature of problem:*

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44 The program calculates the poles of the partial wave S-matrix for spherically sym-
45 metric strictly finite range complex potentials. A few types of potential forms are
46 built in and option for reading in external potential form is given. Landscape of
47 the S-matrix on a given domain of the complex wave number plane can be cal-
48 culated. Accurate position of the poles can be determined. Normalized Gamow
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54 ¹The program name is chosen to honor the late József Zimányi to whom one of the
55 authors (T. Vertse) is grateful for starting his carrier.
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59 *Preprint submitted to Computer Physics Communications*

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9 wave functions and trajectories of given poles can be calculated optionally.

10 *Solution method:*

11 Internal and external solution satisfying boundary conditions in the origin and in
12 the asymptotic region are generated by integrating the radial equation with adap-
13 tive step-size control for Runge-Kutta method. The difference of the logarithmic
14 derivatives are calculated for a range of distances. The minimum of the summed
15 modulus of the differences is searched using the Nelder-Mead algorithm. Pole tra-
16 jectories and normalized Gamow functions can be calculated optionally.

17 *Additional comments including Restrictions and Unusual features:*

18 The region of interest is restricted to the lower half of the wave number plane.
19 Pole solutions from the upper half wave number plane can be safely computed by
20 using the codes GAMOW [1] and ANTI [2].
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26 **References**

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28 [1] T. Vertse, K. F. Pál, Z. Balogh, Computer Physics Communications **27**, 309
29 (1982).
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31 [2] L.Gr. Ixaru, M. Rizea, T. Vertse, Computer Physics Communications **85**, 217
32 (1995).
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36 **1. Introduction**

37
38 One of the possibility of defining resonances is by the purely outgoing
39 solutions of the Schroedinger equation (Siegert condition), i.e. by solutions
40 with complex energies at the pole of the S -matrix. These solutions are called
41 Gamow states, since Gamow was the first who introduced them in nuclear
42 physics for the description of α -decay early in the last century. Gamow
43 states represent non-stationary states, since they describe resonant states
44 with finite lifetime. The lifetime is promotional to the inverse of the width of
45 the resonance. The energy of the Gamow solution is a discrete complex value
46 which corresponds to the pole of the S -matrix in the complex energy sheets.
47 It is more convenient to use the wave number k instead of the energy, since in
48 single channel problem we have only one complex wave number sheet. The
49 energy is proportional to the square of k , therefore the upper half of the k -
50 plane maps to one of the energy sheets (called physical energy sheet), while
51 the lower half of the k -plane maps to another energy sheet, called unphysical
52 energy sheet.
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Decaying resonances lie in the fourth quadrant of the k -plane. For a real potential the capturing resonances are mirror images of the decaying resonances. Both type of resonances lie on the second energy sheet, the decaying ones are below, the capturing ones are above the real axis, where the cut separates the first and the second energy sheets. In the calculation of nuclear reactions we often use complex potential (optical potential). The JOZSO program is able to calculate resonances in complex potentials like its ancestor, the code GAMOW. The name of the new program comes from the nickname of the late József Zimányi to whom one of the authors (T.V.) is greatly indebted.

One of the goal of writing a new code is to provide an efficient tool for calculation the poles of the S -matrix under a whole domain of the complex k plane including regions with large values of the complex k -values. First a landscape of the $S(k)$ -matrix on a mesh of certain region of the k plane is calculated, then the positions of the poles in the domain are localized. After we localized a complex k eigenvalue we normalize its wave function to unity, using regularization methods taking the contribution of the external region into account. For neutral particle the contribution of the external region can be calculated in closed form given in Ref.[9]. It turned out only later that the normalized wave function of the anti-bound pole can be either real or imaginary in a real potential [11]. An option for calculating trajectory of selected poles is built into the new program. The trajectories are complex curves in the complex wave number plane along a certain pole moves as the potential strength slowly changes. The potential strength is a real number γ , which multiplies the sum of the nuclear potential terms. The $\gamma \rightarrow 0$ represents the free particle limit. The use of the code JOZSO was demonstrated in Refs. [2] and [4]. In a most recent work [12] it was found that the JOZSO program gives more accurate values for the imaginary part of the resonance for large values of the real part of k .

The program deals with potential wells only, i.e. in which the poles become bound or anti-bound for large values of the potential depth.

We describe the mathematical formalism in details in the following sections.

2. Formalism

The complex energy of the Gamow solution is $E = E^R - iE^I$, where E^R is the position of the resonance and $E^I = \Gamma/2$ with the width Γ . The

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9 wave number of the Gamow resonance is also complex, $k = k^R + ik^I$. Since
10 k^I is negative, therefore the complex k lies on the lower half of the com-
11 plex k -plane. The real part k^R is positive for a decaying resonance, while a
12 resonance with negative real part k^R corresponds to a capturing resonance.
13 The energy is proportional to the square of k , therefore the resonant energy
14 is on the second Riemann-sheet, it is a discrete complex eigenvalue of the
15 differential equation. Strictly speaking it is a generalized eigenvalue since
16 E is not real as an eigenvalue in normal sense should be. We assume that
17 the phenomenological potential is spherically symmetric, therefore the use of
18 the polar coordinates is convenient. The resonant wave function satisfies the
19 radial Schroedinger equation as follows:
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$$23 \quad u''(r, k) + \left[k^2 - \frac{l(l+1)}{r^2} - v(r) \right] u(r, k) = 0 , \quad (1)$$

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27 where prime denotes the derivative with respect to the radial distance r . The
28 non-negative integer l denotes the quantum number of the orbital angular
29 momentum, $v(r)$ denotes the sum of the nuclear and Coulomb potentials
30 both having spherical symmetry. In our case we have no Coulomb potential
31 term. We can rewrite the radial equation in Eq.(1) into the form expressed
32 by the so called squared local wave number:
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$$36 \quad k_l^2(r) = \left[k^2 - \frac{l(l+1)}{r^2} - v(r) \right] , \quad (2)$$

$$37 \quad u''(r, k) + k_l^2(r)u(r, k) = 0 . \quad (3)$$

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42 The solution $u(r, k)$ satisfies boundary condition (BC) at the origin $r = 0$
43 and at large distance $r \geq R_{as}$, beyond the range of the nuclear potential
44 R_{max} . The energy E and the potentials too are written in the same units as
45 k^2 and the centrifugal term: $\frac{l(l+1)}{r^2}$, namely in $[fm^{-2}]$. The factor $c_1 = \frac{2\mu}{\hbar^2}$
46 converts from the usual MeV units to $[fm^{-2}]$ and it includes the reduced
47 mass $\mu = \frac{m_p m_T}{(m_p + m_T)}$ of the projectile-target system. Therefore
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$$51 \quad k^2 = c_1 E \quad v(r) = c_1 V(r) , \quad (4)$$

52
53 where $V(r)$ denotes the total nuclear potential. The nuclear potential we
54 deal with here has a strictly finite range (SFR) feature, i.e it falls to zero and
55 remains zero beyond a finite distance.
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9 The BC at the origin requires that the solution of the radial equation is
10 regular:

$$11 \quad u(0, k) = 0 . \quad (5)$$

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13 The other BC is specified in a large distance: $R_{as} \geq R_{max}$, where the nuclear
14 potential vanishes:

$$15 \quad V(r \geq R_{max}) = 0 . \quad (6)$$

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18 *2.1. Asymptotic forms of the SFR potentials and the solutions*

19 At R_{as} i.e. at or beyond R_{max} our radial equation in Eq.(1) evolves to its
20 asymptotic form without potential. It describes free spherical waves, which
21 satisfy the Riccati–Hankel differential equation

$$22 \quad u''(r, k) + \left[k^2 - \frac{l(l+1)}{r^2} \right] u(r, k) = 0 . \quad (7)$$

23
24 It is convenient to change to the dimensionless variable $\rho = kr$ in the asymp-
25 totic differential equation.

26
27 For a scattering state the asymptotic BC requires that the solution $u(r, k)$
28 should be a linear combination of the incoming $I_l(kr)$ and outgoing $O_l(kr)$
29 spherical free waves:

$$30 \quad u(r, k) = A[I_l(kr) - S(k)O_l(kr)] , \quad (8)$$

31
32 where $S(k)$ is the element of the scattering matrix. In this case the scattering
33 matrix is diagonal in the angular momentum, therefore in the partial wave l
34 $S(k)$ is a 1×1 matrix, the scattering function. The incoming $I_l(kr)$ and out-
35 going $O_l(kr)$ spherical free waves are expressed by Ricatti-Hankel functions
36 $H_l^\pm(\rho)$. Here $H_l^+(\rho) \sim O_l(\rho)$ and $H_l^-(\rho) \sim I_l(\rho)$.

37
38 Solutions being regular at the origin at any real or complex k values can
39 be matched at R_{as} to the combinations of $I_l(kr)$ and $O_l(kr)$ solutions of the
40 asymptotic differential equation.

41
42 For the majority of the potential forms used in nuclear calculations the
43 radial equation in Eq. (1) can not be solved analytically and the solution
44 $u(r, k)$ should be calculated by using numerical solution methods with dif-
45 ferent approximations. At $r = R_{as}$ the numerical solution should match to
46 that of the asymptotic equation in Eq.(8). The derivative $u'(r, k)$, what we
47 also calculate numerically should be equal to the derivative of the asymptotic
48 equation, therefore

$$49 \quad u'(r, k) = A k [\bar{I}_l(kr) - S(k)\bar{O}_l(kr)] , \quad (9)$$

where $\bar{I}_l(kr)$ and $\bar{O}_l(kr)$ denote the derivatives of the incoming and outgoing Ricatti-Hankel functions with respect to $\rho = kr$. The magnitude A of the asymptotic solution falls out from the logarithmic derivative if r is in the asymptotic region. We can calculate the value of $S(k)$ from the logarithmic derivative at $r = R_{as}$:

$$z_i(R_{as}, k) = \frac{u'_i(R_{as}, k)}{u_i(R_{as}, k)} = k \frac{\bar{I}_l(kR_{as}) - S(k)\bar{O}_l(kR_{as})}{I_l(kR_{as}) - S(k)O_l(kR_{as})}. \quad (10)$$

Here we denote by $u_i(r, k)$ the internal solution being regular at $r = 0$.

The value of the S -matrix at the wave number k can be calculated from this relation:

$$S(k) = \frac{k\bar{I}_l(kR_{as}) - z_i(R_{as}, k)I_l(kR_{as})}{k\bar{O}_l(kR_{as}) - z_i(R_{as}, k)O_l(kR_{as})}. \quad (11)$$

If we solve the differential equation in Eq.(1) numerically, we get at $r = R_{as}$ the logarithmic derivative $z_i(R_{as}, k)$ needed for calculating the value of $S(k)$ in Eq.(11). In the nuclear reaction calculations we need the values of the $S(k)$ in each partial waves to calculate cross sections.

For a resonance the asymptotic BC requires that the solution $u(r, k)$ should be proportional to the outgoing function $O_l(kr)$ and its derivative $u'(r, k)$ should be proportional to the derivative of the outgoing function, i.e.

$$u(r, k) \sim O_l(kr) \quad u'(r, k) \sim k\bar{O}_l(kr), \quad (12)$$

or the logarithmic derivative should be

$$z(r, k) = k \frac{\bar{O}_l(kr)}{O_l(kr)}. \quad (13)$$

This BC can be satisfied only at a discrete complex k_o eigenvalue, namely at a k -value belonging to a pole of $S(k)$. The procedure of finding the poles of $S(k)$ will be discussed in the section 2.6.

Since the position of the pole depends on the potential (on its strength and on its shape) we can calculate the pole trajectory for a potential with given radial shape by calculating the pole position k_o as a function of the potential strength. By increasing the strength the trajectory goes to the upper half of the k -plane, where (for real potential) it becomes a bound state with real energy and with purely imaginary k . Here the solution becomes to be a square integrable real function with finite number of nodes n . We assign the $n = 0$

node number to the function having zero only at $r = 0$. The $n = 1$ solution has one additional node at $r > 0$, the $n = 2$ solution has two additional nodes at $r > 0$, etc. Anti-bound solutions are not square integrable since they diverge as $r \rightarrow \infty$. After normalizing them with appropriate regularization procedure they become functions being either real or purely imaginary [11] and they have finite number of nodes. Resonances on the other hand have no nodes, only infinite number of zeros both in the real part and in the imaginary part of their wave function at different distances. The asymptotic solution $O_l(kr)$ oscillates around the real r -axis with exponentially growing amplitude as $r \rightarrow \infty$. To assign a given node number to a resonance is possible only if we manage to get rid off the oscillations in the asymptotic region by following the pole trajectory until it becomes a bound state with finite node number.

2.2. Nuclear potentials

A common feature of the nuclear potentials is that they differ from zero only in a finite range $r \in [0, R_{max}]$. Their functional shapes (radial forms) can be built into this program, or might be read in from file (external form factor). The built in potentials might have different phenomenological forms. A common feature of these potentials is that the radial equation in Eq.(1) is solved numerically. The most frequently used potential is the Woods-Saxon (WS) potential. Most of the $V(r)$ potentials, including the WS becomes zero only at infinite distance. However, the nuclear potential $V(r)$ we use here, should be SFR type potential, since we match our solution at R_{as} to that of the asymptotic differential equation in which no nuclear part is present[11]. The most common nuclear potential is the cut-off form of the Woods-Saxon (CWS) potential. The CWS can be written as a product of its strength V_0 and its radial shape:

$$V^{CWS}(r, R, a, R_{max}) = -V_0 f_{CWS}(r, R, a, R_{max}) , \quad (14)$$

where the radial shape is

$$f_{CWS}(r, R, a, R_{max}) = \theta(R_{max} - r) \frac{1}{1 + e^{\frac{r-R}{a}}} , \quad (15)$$

where $\theta(x)$ denotes the Heaviside step function, being zero for negative and unity for non-negative arguments. It was shown earlier [7] that in the CWS potential the positions of broad resonances considered here do depend on the

value of the cut-off radius R_{\max} [13, 11], therefore, the cut-off radius is an important parameter of the CWS form in Eq. (15). The two other parameters of the CWS form are the radius R and the diffuseness a .

The generalized WS potential (GWS) is a combination of a Woods-Saxon (WS) potential term and a surface term with potential strengths V_0 and V_1 . The radial form of the WS term is

$$f^{\text{WS}}(r, R, a) = - \frac{1}{1 + e^{\frac{r-R}{a}}} , \quad (16)$$

while the shape of the surface term is

$$f^{\text{SWS}}(r, R, a) = - \frac{e^{\frac{r-R}{a}}}{(1 + e^{\frac{r-R}{a}})^2} . \quad (17)$$

The geometrical parameters of the terms are the radius R and diffuseness a . Therefore the resulting GWS potential is the following:

$$V^{\text{GWS}}(r, R, a, V_0, V_1) = V_0 f^{\text{WS}}(r, R, a) + V_1 f^{\text{SWS}}(r, R, a) . \quad (18)$$

A big advantage of this potential is, that for $l = 0$ the radial equation can be solved in closed analytic form [3]. A SFR form of the GWS potential can be created, if we cut its radial form at the same finite distance R_{\max} as we do it with the radial form of its volume term. This SFR form $f^{\text{CGWS}}(r)$ has the property of being zero at and beyond the distance R_{\max} . While the analytical solution exists only if the geometrical parameters R, a are the same for the volume and the surface terms, for a numerical solution this is not needed and we can use different parameters for the two terms.

A new type of phenomenological nuclear potential form (SV form) was introduced in Ref. [7]. The SV form becomes zero smoothly at a finite R_ρ distance without an artificial cut-off and remains zero beyond that distance. The SV form has the attractive mathematical property that it belongs to the class of functions C^∞ , the functions of compact support as it was realized by Nándori[10].

Here we specify the SV potential form as a product of its strength and radial shape

$$V^{\text{SV}}(r) = -V_0 f^{\text{SV}}(r, c, \rho_0, \rho_1) , \quad (19)$$

in which the strength $V_0 \geq 0$, and the shape $f^{\text{SV}}(r, c, \rho_0, \rho_1)$ is a linear combination of the function

$$f(r, \rho) = e^{\frac{r^2}{r^2 - \rho^2}} \theta(\rho - r) , \quad (20)$$

and a term containing the derivative, with respect to r , of the first factor,

$$f'(r, \rho) = -\frac{2r\rho^2}{(r^2 - \rho^2)^2} e^{\frac{r^2}{r^2 - \rho^2}} \theta(\rho - r) . \quad (21)$$

$$f^{\text{SV}}(r, c, \rho_0, \rho_1) = f(r, \rho_0) - cf'(r, \rho_1) . \quad (22)$$

The combination parameter c gives the weight of the derivative term. For light nuclei the derivative term is not important and one can take $c = 0$ [8].

Sahu and Sahu [5] generalized the SV potential by introducing an extra parameter a_s to the derivative term of the SV form. The formula of the SS potential [5] is analogous to Eq. (22), where the SS form is given as:

$$f^{\text{SS}}(r, c, \rho_0, \rho_1, a_s) = f(r, \rho_0) - cf'(r, \rho_1, a_s) , \quad (23)$$

where

$$f'(r, \rho_1, a_s) = -\frac{2r\rho_1^2}{(r^2 - \rho_1^2)^2} e^{\frac{a_s r^2}{r^2 - \rho_1^2}} \theta(\rho_1 - r) , \quad (24)$$

with the extra diffuseness parameter a_s . When $a_s = 1$, the SS form coincides with the SV potential (19). By using $a_s \neq 1$, one naturally has more freedom in choosing the shape of the potential. With the usual choice $\rho_0 > \rho_1$, the range of the SS potential is also ρ_0 . The SS form has the same attractive mathematical features as the SV potential, namely it is a C^∞ function.

The external form factors are read in from the input file at specified points. Since we need the potentials later at points not specified in advance, it is convenient to interpolate the external forms using spline interpolations. If we have the spline coefficients of the potentials we are able to calculate them at any point inside the range they are different from zero.

If we are calculating pole trajectories we can determine the spline coefficients only once and multiple the interpolated shape with the potential strength we vary.

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9 *2.3. Spin-orbit part of the nuclear potential*

10 Since the neutron has a non-zero spin ($s = 1/2$ in \hbar unit) the potential
11 v can be complemented by a spin-orbit term:
12

$$13 \quad V_{\text{so}}^{\text{CGWS}}(r, R_{\text{so}}, a_{\text{so}}, R_{\text{max}}) = V_{\text{so}}^{\text{CGWS}} h_{\text{CWS}}(r, R_{\text{so}}, a_{\text{so}}, R_{\text{max}}) 2(\mathbf{l} \cdot \mathbf{s}) , \quad (25)$$

14 with a radial form
15

$$16 \quad h_{\text{CGWS}}(r, R, a, R_{\text{max}}) = -\frac{1}{r} f'_{\text{CGWS}}(r, R, a, R_{\text{max}}) , \quad (26)$$

17 in which the derivative of the central potential appears. The spin-orbit term
18 of the SS potential may be defined analogously:
19

$$20 \quad V_{\text{so}}^{\text{SS}}(r, c, \rho_0, \rho_1) = V_{\text{so}}^{\text{SS}} h_{\text{SS}}(r, c, \rho_0, \rho_1) 2(\mathbf{l} \cdot \mathbf{s}) , \quad (27)$$

21 with
22

$$23 \quad h_{\text{SS}}(r, c, \rho_0, \rho_1) = -\frac{1}{r} f'_{\text{SS}}(r, c, \rho_0, \rho_1) = \frac{2\rho_0^2}{(r^2 - \rho_0^2)^2} e^{\frac{r^2}{r^2 - \rho_0^2}} \theta(\rho_0 - r) \quad (28)$$

$$24 \quad - c \frac{2\rho_1^2}{(r^2 - \rho_1^2)^4} e^{\frac{a_s r^2}{r^2 - \rho_1^2}} \left(\frac{\rho_1^4}{r} - 3r^3 + 2r\rho_1^2(1 - a_s) \right) \theta(\rho_1 - r)$$

25 The spin-dependent factor $2(\mathbf{l} \cdot \mathbf{s})$ can be calculated easily from the difference
26 of the eigenvalues of the total, the orbital and the spin quantum numbers:
27

$$28 \quad 2(\mathbf{l} \cdot \mathbf{s}) = j(j + 1) - l(l + 1) - s(s + 1) . \quad (29)$$

29 Due to the $1/r$ factor in the spin-orbit potential, it might be singular at the
30 origin. However, for the SS form with $c = 0$ the singularity disappears. In
31 the general case the full potential is:
32

$$33 \quad V(r) = V_{\text{cent}}(r) + V_{\text{so}}(r) . \quad (30)$$

34 Here the $V_{\text{cent}}(r)$ central part of the nuclear potential does not depend on
35 j , while the spin-orbit part might depend on j if the spin-orbit strength V_{so}
36 is different from zero. In certain cases the central potential might depend
37 on the orbital angular momentum l . The pair of the l, j quantum numbers
38 define a given partial wave of the scattering problem. In a given partial wave
39 the nuclear potential might depend on l and j , i.e. $V^{l,j}(r)$ and the local wave
40 number also might depend on j as well $k_{l,j}^2(r)$. The radial wave function
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$u(r, k)$ depends on the potential used in the given partial wave, but this dependence is not shown explicitly in the notation for the sake of simplicity. The value of the S -matrix in the given partial wave depends on j too, if the spin-orbit part of the potential is different from zero. When we calculate pole trajectory, the potential in Eq. (30) is multiplied by the strength γ and the value of γ is changing along the trajectory.

2.4. Normalization of the wave function

To find the pole position we do not need the normalized wave function. Assume that the pole is at k_o , then the non-normalized resonance solution we have in $u_i(r, k_o)$. However, we can compute the normalized solution choosing WF=1 in the input file.

The normalization of this radial wave function of the Gamow resonance can be done most conveniently by using the method given in Ref.[9]. The same method was implemented in the program GAMOW [1] and also in the program ANTI [6]. The contribution to the square of the norm from the internal region can be calculated by quadrature as:

$$N_i^2 = \int_0^{R_{as}} u_i(r, k_o)^2 dr . \quad (31)$$

This should be complemented by the contribution of the external region given in closed form [9] as

$$\begin{aligned} N_e^2 &= F_c^2 \int_{k_o R_{as}}^{\infty} H_l^+(\rho)^2 d\rho \\ &= \frac{-R_{as} F_c^2}{2} [H_l^+(k_o R_{as})^2 + H_{l+1}^+(k_o R_{as})^2 - \frac{2l+1}{k_o R_{as}} H_l^+(k_o R_{as}) H_{l+1}^+(k_o R_{as})], \end{aligned} \quad (32)$$

where $F_c = \frac{u_i(R_{as}, k_o)}{H_l^+(k_o R_{as})}$. The role of the factor F_c is to continue the internal solution smoothly into the asymptotic region.

Now the full squared norm is as follows

$$N^2 = N_e^2 + N_i^2 , \quad (33)$$

while the normalized radial wave function is:

$$u(r, k) = \frac{1}{N} u_i(r, k) . \quad (34)$$

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9 *2.5. Numerical integration of the radial equation*

10 For the numerical integration of the radial equation we use the adaptive
11 step-size control for Runge-Kutta-Fehlberg with Cash-Karp parameters as in
12 [14] (SUBROUTINE ODEINT with RKQC).
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15 *2.6. Finding the poles of $S(k)$.*

16 We have more chance for finding the pole if we can start the iterations
17 with good k starting value. In this case the calculation is very efficient, since
18 we are solving a set of initial value problems instead of eigenvalue problem
19 as follows.
20

21 In the eigenvalue problem a possibility for finding the pole of S is that
22 we calculate the external solution with outgoing wave starting condition at
23 R_{as} and from these initial values at R_{as} we propagate the numerical solution
24 inward and calculate the external solution $u_e(r_j, k)$ in the mesh-points r_j
25 needed to compare it to the internal solution.
26
27

28 From the BC at the origin we can start an internal solution of the radial
29 equation in Eq.(1) with the starting values:
30

$$31 \quad u_i(0, k) = 0 \quad u'_i(0, k) = 1 . \quad (35)$$

32 Both BC (in Eqs. (5) and (13)) can be satisfied simultaneously only at
33 discrete complex k eigenvalues belonging to a poles of $S(k)$. Here the complex
34 k eigenvalue is fixed by the zeros of the the difference of the logarithmic
35 derivatives of the internal and the external solutions:
36
37

$$38 \quad G(k, r) = z_i(r, k) - z_e(r, k) , \quad (36)$$

39 where

$$40 \quad z_i = \frac{u'_i(r, k)}{u_i(r, k)} \quad \text{and} \quad z_e = \frac{u'_e(r, k)}{u_e(r, k)} .$$

41 The computer programs GAMOW[1], and ANTI[6] find the zeros of $G(k, r)$,
42 at certain R_m matching radius $0 < R_m < R_{as}$. For a broad resonance the
43 proper choice of this R_m is difficult. The zero is searched by Newton itera-
44 tions, and the iteration process often converges poorly or fails. Therefore in
45 the JOZSO program we extend the comparison of the logarithmic derivatives
46 of z_i and z_e to a wider region of r and we search for the absolute minimum
47 of the following function:
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$$54 \quad F(x_1, x_2) = \log \left[\sum_{j=i_1}^{i_2} |G(k, r_j)| \right] . \quad (37)$$

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9 where r_j is a mesh with equidistant mesh-points of the interval $r = [r_1, r_2]$,
10 with step-length h . Here the two real variables in the argument of F are the
11 real and the imaginary parts of the complex wave number k (or vice versa).
12 The logarithmic derivatives are complex, hence in order to have a real valued
13 function we take the modulus of their differences. Therefore the function F
14 is a real valued function. Although we calculate the logarithmic derivatives
15 of the internal and the external solutions in a wide range of r , in the sum in
16 Eq. (37) we include only a subinterval in which the nuclear potential falls to
17 the certain fraction to their value close to the origin. We take the value of
18 the potential close to the origin, say e.g. $V_{or} = V_N(h)$ and select the index i_1
19 where $V(i_1 * h = r_1) \approx V_{or}/10$. The higher value of the subscript is taken as
20 $V(i_2 * h = r_2) \approx V_{or}/1000$.
21
22

23
24 The task is to find the absolute minimum of the real function F . If the
25 two solutions could be calculated without any errors than the value of the
26 absolute minimum of the sum in the argument of the logarithm were zero.
27 But in the numerical solution rounding errors accumulate as we proceed from
28 the starting points, where the initial conditions are specified, therefore this
29 value is somewhat larger than zero. The minimum of the function $F(x_1, x_2)$
30 in Eq.(37) should have a negative value with a large modulus.
31
32

33 To find the minimum of the function $F(x_1, x_2)$ we use the Nelder–Mead
34 method. We are searching for a minimum of the function starting from some
35 first guess k_0 . The two-variable functions $F(x_1, x_2)$ in Eq.(37) have special
36 shapes, therefore it is useful to generate a landscape of the $F(x_1, x_2)$ over a
37 grid of a region of the complex k domain. The mesh in x_1 and x_2 should be
38 fine enough to localize the poles of $S(k)$ and at the same time the number
39 of mesh-points should remains within a reasonable limit. The landscape
40 helps us to supply reasonably good starting values for finding the absolute
41 minima of the function $F(x_1, x_2)$ and give the position of the pole with higher
42 accuracy. Pole positions calculated agree with those given in Ref.[6] in 3-4
43 decimal digits.
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48 *2.7. Starting integration from the origin*

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50 The centrifugal term for $l > 0$ has a singularity in Eq.(2), therefore it
51 might be convenient from the numerical point of view to use an expansion
52 method for the regular solution close to the origin. Here (at the first *nbo*
53 points) we can expand the internal solution $u_i(r, k)$ into powers of r and we
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search for the solution for $r \ll 1$ in a form:

$$u_i(r, k) = \sum_{j=0}^{50} a_j r^{l+1+j} . \quad (38)$$

The full potential in Eq.(6) can be split into two terms, the one which is singular at the origin ($v_{sg}(r)$) and the rest which is not singular ($v_{ns}(r)$). For a neutron the only term which might be singular at $r = 0$ is the spin-orbit potential $v_{so}(r)$ with radial form in Eq.(26). The spin-orbit potential is non-zero only for $l > 0$ and for non-zero spin-orbit strength, in this case it can be approximated as b/r . For the SS potential the second term (Eq.(24)) is proportional to r , therefore $b = 0$. Let us approximate the non-singular potential with a parabola and write the full potential for $r \ll 1$ as

$$v(r) \approx b/r + v_0 + v_1 r + v_2 r^2 . \quad (39)$$

With this the local wave number in Eq.(2) can be approximated as

$$k_{l,j}^2(r) = a - v_1 r - v_2 r^2 - \frac{l(l+1)}{r^2} - \frac{b}{r} . \quad (40)$$

The coefficients of the expansion in Eq.(38) can be calculated as:

$$a_0 = 1 \quad a_1 = \frac{a_0 b}{2l+2} \quad a_2 = \frac{a_1 b - a_0 a}{4l+6} \quad a_3 = \frac{a_2 b - a_1 a + a_0 v_1}{6l+12}$$

and for $i > 3$ the following recurrence relation holds

$$a_i = \frac{1}{i(2l+i+1)} [a_{i-1} b - a_{i-2} a + a_{i-3} v_1 + a_{i-4} v_2] . \quad (41)$$

The derivation of the form in Eq.(38) is straightforward, and it gives $u'(r, k)$ at the first equidistant mesh-points: $r_j = j * h$, for $j = 1, \dots, nbo$, and the logarithmic derivatives will be

$$z_i(r, k) = \frac{\sum_{j=0} a_j (l+j+1) r^{l+j}}{\sum_{j=0} a_j r^{l+1+j}} . \quad (42)$$

Having the solution and its derivative at $r_{nbo} = nbo * h$ we can proceed for $r > r_{nbo}$ and calculate the solution using the integration routine DIFFSOLVE with the required accuracy *eps*. This way we propagate the internal solution $u_i(r, k)$ or the logarithmic derivative function $z_i(r, k)$ from this r_{nbo} point outward, until we reach the asymptotic region at $r = R_{as}$.

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9 *2.8. Calculation of the asymptotic solutions*

10 For neutron the asymptotic solution goes to the differential equation of
11 the Ricatti-Hankel functions $H_l^\pm(\rho)$. Here $H_l^+(\rho) \sim O_l(\rho)$ and $H_l^-(\rho) \sim I_l(\rho)$
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13

$$14 \frac{d^2 w(\rho)}{d\rho^2} + \left[1 - \frac{l(l+1)}{\rho^2}\right] w(\rho) = 0. \quad (43)$$

15
16 The Ricatti-Hankel functions can be calculated easily by using the three
17 terms recurrence relation:
18
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$$20 H_{l+1}^\pm(\rho) = \frac{2l+1}{\rho} H_l^\pm(\rho) - H_{l-1}^\pm(\rho). \quad (44)$$

21 To start the recurrence we use the known form of the $l = 0, 1$ functions:
22
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$$24 H_0^\pm(\rho) = e^{\pm i\rho} \quad H_1^\pm(\rho) = \left(\frac{1}{\rho} - i\right) e^{\pm i\rho}. \quad (45)$$

25 For $l = 0$ the derivative of the Ricatti-Hankel function is simply $\bar{H}_0^\pm(\rho) =$
26 $\pm i H_0^\pm(\rho)$. For $l > 0$ the derivatives can be calculated from the relation:
27
28

$$29 \bar{H}_l^\pm(\rho) = H_{l-1}^\pm(\rho) - \frac{l}{\rho} H_l^\pm(\rho). \quad (46)$$

30 For a Gamow resonance the external BC fixes the external solution of the
31 radial equation as:
32
33

$$34 u_e(r = R_{as}, k) = O_l(kR_{as}) = H_l^+(kR_{as}), \quad (47)$$

$$35 u'_e(r = R_{as}, k) = k\bar{O}_l(kR_{as}) = k\bar{H}_l^+(kR_{as}),$$

36 with the actual value of the complex wave number k . We can propagate the
37 numerical solution from this point inward and calculate the external solution
38 $u_e(R_{as}, k)$.
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43 *2.9. Trajectory calculation*

44 For the calculation of the trajectory of a given pole, we can introduce a
45 real parameter γ for the nuclear potential and define the nuclear potential as
46 a product of the full nuclear potential and this strength γ . Then we consider
47 the pole position as a function of the real variable γ .
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49 We start from a strength $\gamma = 1$ and increase or decrease the strength
50 in certain $\delta\gamma$ steps. The number of steps in γ is limited by an input data
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9 of the program. The sign of the step $\delta\gamma$ defines if we go up or down along
10 the trajectory. A positive step $\delta\gamma > 0$ increases the γ and we use deeper
11 and deeper potentials. The negative value has the opposite effect and we
12 proceed downward along the trajectory toward the starting (small) value of
13 the potential. The move of the poles along their trajectories was studied
14 recently in Ref. [4].

15
16 If we have the pole wave number of the resonance k_0 at a potential
17 strength γ we can start to calculate pole trajectory $k_0(\gamma)$ starting from this
18 point. We modify the potential strength with a small $\delta\gamma$ value and calcu-
19 late the pole with this modified value to get $k_0(\gamma + \delta\gamma)$. If the change $\delta\gamma$ is
20 small then the change of the $\delta k_0 = k_0(\gamma + \delta\gamma) - k_0(\gamma)$ is also small and the
21 convergence to the pole of the changed potential will be fast and reliable.
22

23
24 We can calculate the pole trajectories from their starting points where γ is
25 very small and we can proceed by increasing their strength. The advantage of
26 using this tactic is that the starting values of the trajectories show regularity
27 for the strictly finite range potentials as has been discussed in Ref.[8]. The
28 shape of the trajectory naturally depends on the radial shape of the potential
29 and also on the partial wave l, j and the sequence number m of the pole. The
30 index m increases as the k^R increases, the $m = 1$ is the pole being closest
31 to the imaginary k -axis. The index m is different from the node number n
32 what we can determine only when we continue the trajectory to the upper
33 half of the k -plane and we have a bound state pole with definite number of
34 nodes. The larger is the index m the farther is the starting point k_s from the
35 imaginary k -axis, and we have to follow the pole along a long way until the
36 resonance becomes a bound state in a deep potential.
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39 We can follow another tactics and start the trajectory from the bound
40 state in a deep potential, where we can count its node easily. Then we make
41 the potential shallower by taking $\delta\gamma < 0$ and proceed along the trajectory
42 until we cross the origin $k = 0$ and go to the lower half of the k -plane. For
43 $l > 0$ where we have a centrifugal barrier the bound state pole goes to a
44 resonance, while for $l = 0$ (no barrier) the bound state first becomes to be
45 anti-bound state and might become resonance when the pole departs from
46 the imaginary k -axis. This situation is studied in Ref.[11] extensively.
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52 *2.10. The use of external potential form*

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54 Sometimes it is useful to use potential form calculated separately, by
55 using different theoretical models. In this case the geometrical form can be
56 given in a file *extformf.dat* at equidistant h steps in r . After reading in the
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9 external radial form, we use spline interpolation based on the read in knots
10 in order to be able to calculate the external form at any distance needed
11 by the numerical integration routine `diffsolve`. The complex strength of the
12 potential is given in $P(18)+i P(19)$ as with other potential form. A limitation
13 of the present program is that there is no spin-orbit term in the external form.
14 If one wants to include spin-orbit potential one has to add it to the external
15 potential form to be read in.
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19 **3. Program structure**

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21 Input data are read from the file `cgws.config` or `sahu.config` or `ext.config`.
22 There are three possible types of potentials, the type of the potential can be
23 controlled by a command line parameter.

24
25 `jozso -1` or `jozso -cgws` uses CGWS potential,

26
27 `jozso -2` or `jozso -sahu` uses SS potential,

28
29 `jozso -4` or `jozso -ext` uses an external potential,

30
31 the default potential type is the CGWS.

32 The program can be used for calculation a single pole (starting from an initial
33 guess), or for calculation all the poles on a given domain, or for calculation
34 a pole trajectory starting from a given pole. These options are controlled by
35 the `MODE` parameter.

36 The parameters read in are stored in array $P(1), \dots, P(45)$.

37 Meaning of the parameters similar to that in the program `GAMOW`, but
38 some of the parameters are not used or have different role.

39 Here we list the meaning of the input parameters:

40 $P(1)-P(7)$ have the same meaning as in `GAMOW`, namely:

41 $P(1)=AT$ mass of the target,

42 $P(2)=ZT$, charge number of the target (not used),

43 $P(3)=AP$ mass of the projectile,

44 $P(4)=ZP$ charge number of the projectile (not used),

45 $P(5)=LP$ orbital angular momentum of the projectile,

46 $P(6)=JP$ total angular momentum of the projectile,

47 $P(7)=SP$ spin of the projectile,
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53 $P(8)=NBO$ number of points where series expansion is used.

54 Parameters $P(9)-P(10)$ are the initial value of the wave number for `MODE=0`
55 and `MODE=2`. They are not used for `MODE=1`.
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9 P(9)=KR is the real part of the starting value of the complex k .
10 P(10)=KI is the imaginary part of the starting value of the complex k .
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14 Parameters P(11)-P(16) are used for MODE=1 only.
15 P(11)=MESHKR number of mesh-points for the real part of k
16 P(12)=MESHKI number of mesh-points for the imaginary part of k
17 P(13)=DOMAINKR1 lower limit for the real part of k
18 P(14)=DOMAINKI1 upper limit for the imaginary part of k
19 P(15)=DOMAINKR2 upper limit for the real part of k
20 P(16)=DOMAINKI2 lower limit for the imaginary part of k
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25 The parameters P(17)-P(31) are the potential parameters.
26 P(17)=NX gives the type of the nuclear potential.
27 The value of NX is given not in the input file, it is determined from the name
28 of the input file.
29 NX=1 denotes the CGWS potential
30 NX=2 denotes the SS.
31 NX=4 denotes the external potential to be read in from the file extformf.dat.
32 The GWS and SS nuclear potentials are the sum of the central and spin-
33 orbit potential terms. The sums are multiplied by the strength γ , when pole
34 trajectory is calculated.
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38 Parameters of the central potential if NX=1:
39 P(18)=VR is the real part of the volume term of the CGWS potential,
40 P(19)=VI is the imaginary part of the volume term of the CGWS potential,
41 P(20)=R0 is the radius parameter of the volume term, the radius is $R0 \cdot AT^{1/3}$
42 P(21)=A is the diffuseness of the volume term,
43 P(22)=VR1 is the real part of the surface term,
44 P(23)=VI1 is the imaginary part of the surface term,
45 P(24)=R01 is the radius parameter of the surface term, the radius is $R01 \cdot AT^{1/3}$
46 P(25)=A1 is the diffuseness of the surface term,
47 P(26)=GAMMA0 the starting value of the γ parameter used for calculation
48 of pole trajectories.
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54 Parameters of the central potential if NX=2:
55 P(18)=VR is the real part of the volume term of the SS potential,
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 9 P(19)=VI is the imaginary part of the volume term of the SS potential,
 10 P(20)=RHO0 is the range ρ_0 of the first term of the SS potential, it must be
 11 equal to R_{max} in P(38).
 12 P(21)=RHO1 is the range ρ_1 of the second term of the SS potential, it has
 13 to be smaller than ρ_0 ,
 14 P(22)=CS is the relative weight of the second term of the SS potential wrt.
 15 the first term, c in Eqs. (22) and (23).
 16 P(23)=AS is the diffuseness parameter a_s in the second term of the SS po-
 17 tential. The $a_s = 1$ means SV potential,
 18 P(26)=GAMMA0 the starting value of the γ parameter used for calculation
 19 of the pole trajectories.
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25 P(27)=NXSO

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 27 If NXSO=3 then the parameters of the spin-orbit potential for NX=1 are
 28 the following:

29 P(28)=VSOR is the real part of the spin-orbit strength,
 30 P(29)=VSOI is the imaginary part of the spin-orbit strength,
 31 P(30)=R0SO is the radius parameter of the spin-orbit potential for CGWS,
 32 the radius is $R0SO \cdot AT^{1/3}$
 33 P(31)=ASO is the diffuseness parameter of the spin-orbit potential for CGWS.
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38 If NXSO=3 then the parameters of the spin-orbit potential for NX=2 are
 39 the following:

40 P(28)=VSOR is the real part of the spin-orbit strength,
 41 P(29)=VSOI is the imaginary part of the spin-orbit strength,
 42
 43
 44

45 For NX=4 there is only the complex strength in $P(19)+iP(19)$ and the
 46 real strength is:

47 P(26)=GAMMA0 (the starting value of the γ parameter used for calculation
 48 of the pole trajectories).
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52 P(37)=H step size in r , where the radial wave function is calculated.

53 P(38)=RMAX The nuclear potential vanishes beyond this distance. For SS
 54 potential it should be equal to RHO0.

55 P(39)=RAS it is the distance R_{as} where the asymptotic solution is matched
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9 to the numerical one. RAS should be greater than RMAX.
10 P(40)=DGAMMA is the step size of the change of the potential strength γ .
11 P(41)=NGAMMA is the maximal number of changes of the potential strength
12 γ .
13 P(42)=C1 (conversion factor between the energy and the k^2)
14 P(43)=THRESHOLD threshold value for $-F(k)$. Only $-F > P(43)$ peaks
15 are selected as poles.
16 P(44)=MODE
17 For MODE=0 the program searches for a single pole starting from the initial
18 value given by KR and KI.
19 For MODE=1 the program calculates the values of F at the mesh points
20 of the rectangle given by DOMAINKR1, DOMAINKI1, DOMAINKR2, DO-
21 MAINKI2, MESHKR, MESHKI. The mesh points and the function values
22 are written to the file map.dat. The local minima of the function F are de-
23 termined and written to the file peaks.dat. These points are used as starting
24 points for the pole calculation. The poles calculated are written to the file
25 poles.dat.
26 For MODE=2 the program calculates pole trajectory by changing the poten-
27 tial strength by DGAMMA. The change is performed maximum NGAMMA
28 times
29 P(45)=WF calculates the normalized wave function if wf=1 and writes it to
30 the file wf.dat.
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38 Output files in different options.

39 The general output file is jozso.res. Special output files for different options:
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42 For mode=0 the wave number eigen-value k is written to the general out-
43 put file and on the screen as well.
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46 For mode=1 the function values of the F are written to the file map.dat in
47 each mesh points in k . After that the program searches the peaks for which
48 $-F > \text{THRESHOLD}$ and writes them to file peaks.dat. Then it starts search-
49 ing the minima of the function $F(k)$ with starting values of each approximate
50 value. The values of the positions of the minima (poles) are written to file:
51 poles.dat.
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55 For mode=2 the pole trajectory is written to the file ktraj.dat (only the
56 k values) and also to the pltraj.dat in which the potential strength γ is also
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9 written in the first column.

10 11 **Acknowledgment**

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Program file(s)

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